



*2018 Space Symposium*  
**Wavefront Sensing and Controls for the  
James Webb Space Telescope**

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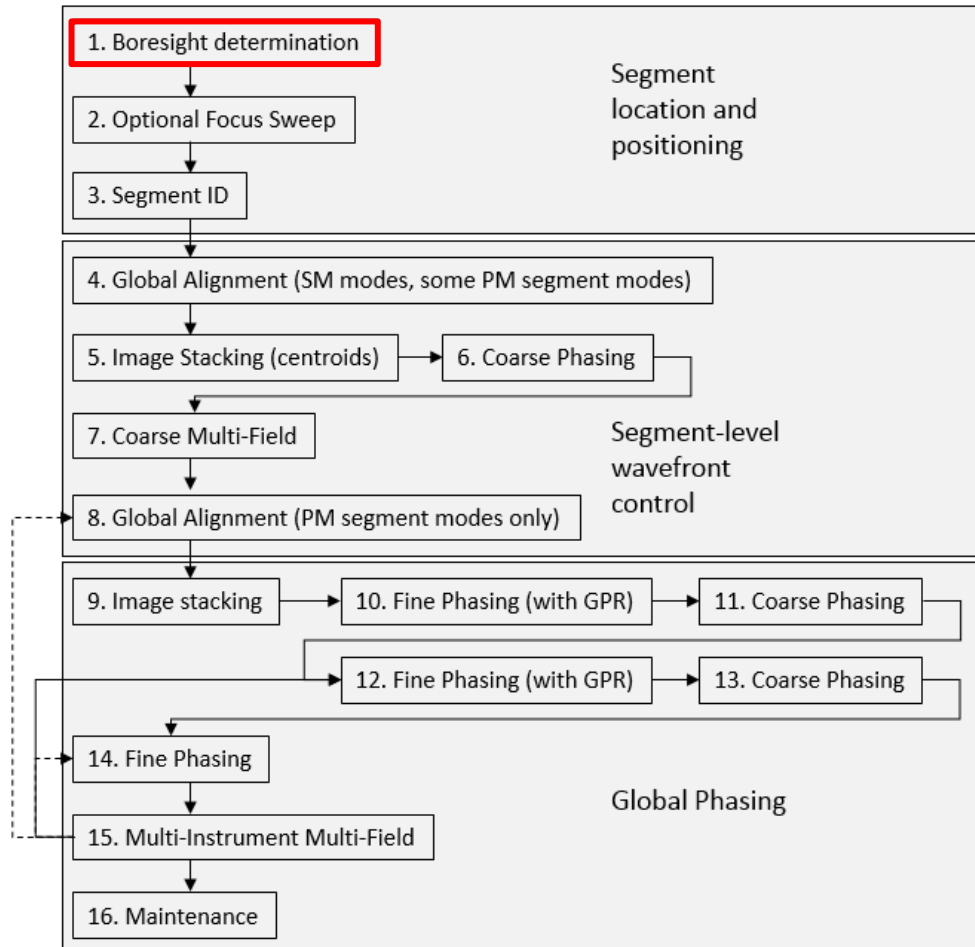


- **James Webb Space Telescope (JWST) is unique in that it is**
  - Large (6.5 meters)
  - Space-based
  - Segmented
  - Deployable
  - Cryogenic
  - High authority
- **The telescope must be aligned in space**
  - Rely on existing science instruments (no wavefront sensor)
  - Special optical elements are placed in 8 of NIRCam's 48 filter wheel positions for WFSC
  - Images are taken and downloaded for analysis on the ground (no on-board image processing)
- **The primary mirror consists of 18 segments**
  - Hexapod allows for adjustment in 6 DOF
  - A 7<sup>th</sup> ROC actuator allows the ROC to be changed
- **The secondary mirror is also on a hexapod**

# The WFSC Process



We will touch on each of these processes briefly



# Image Mosaic and Boresight determination

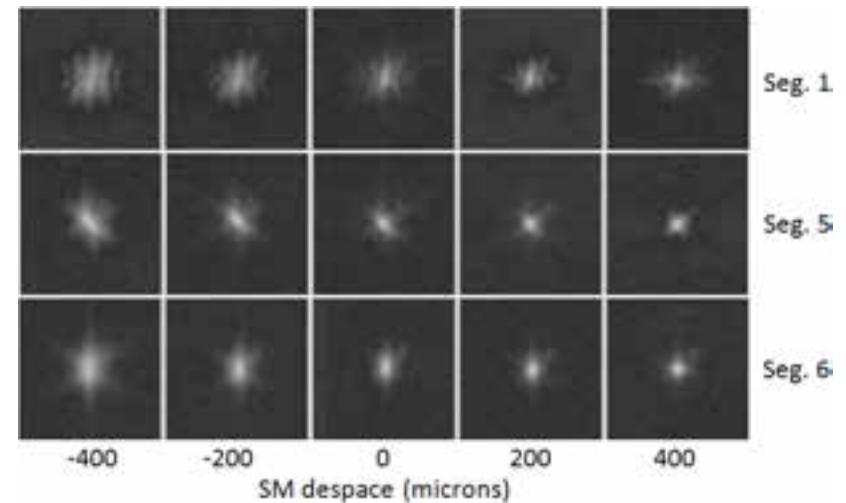


- Telescope is slewed to a bright isolated star.
- ~200 NIRCam images are combined to make a large image mosaic.

# Focus Sweep



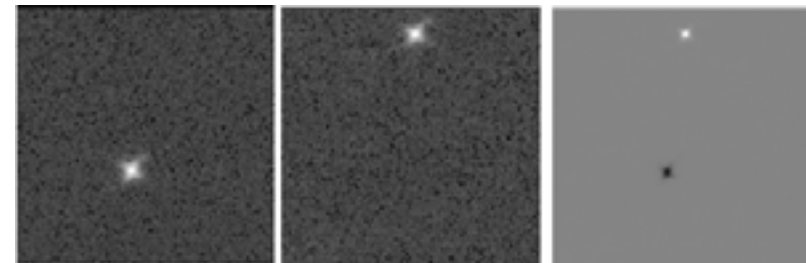
- SM is scanned in piston to step through image focus
- Phase Retrieval calculates focus error
- SM is placed at best focus, while guarantee that at least one spot is focused enough to allow guiding



# Segment ID



- A series of telescope pointings are planned to cover the entire mosaic
- At each pointing, each segment is tilted slightly so its spot will move
- Images are taken before and after tilting. If the image moves, we know which segment it is associated with.



Before tilting

After

Difference



- Spots are moved into a hexagonal array
- SM is moved in piston to create defocused images
- PMSA tilts are also applied to maintain array
- Phase Retrieval yields wavefront over each segment
- Global Alignment is performed twice First application controls
  - SM position
  - PMSA piston
  - PMSA clocking
- Second use of Global Alignment controls
  - PMSA clocking
  - PMSA translation
  - PMSA ROC

# Global Alignment

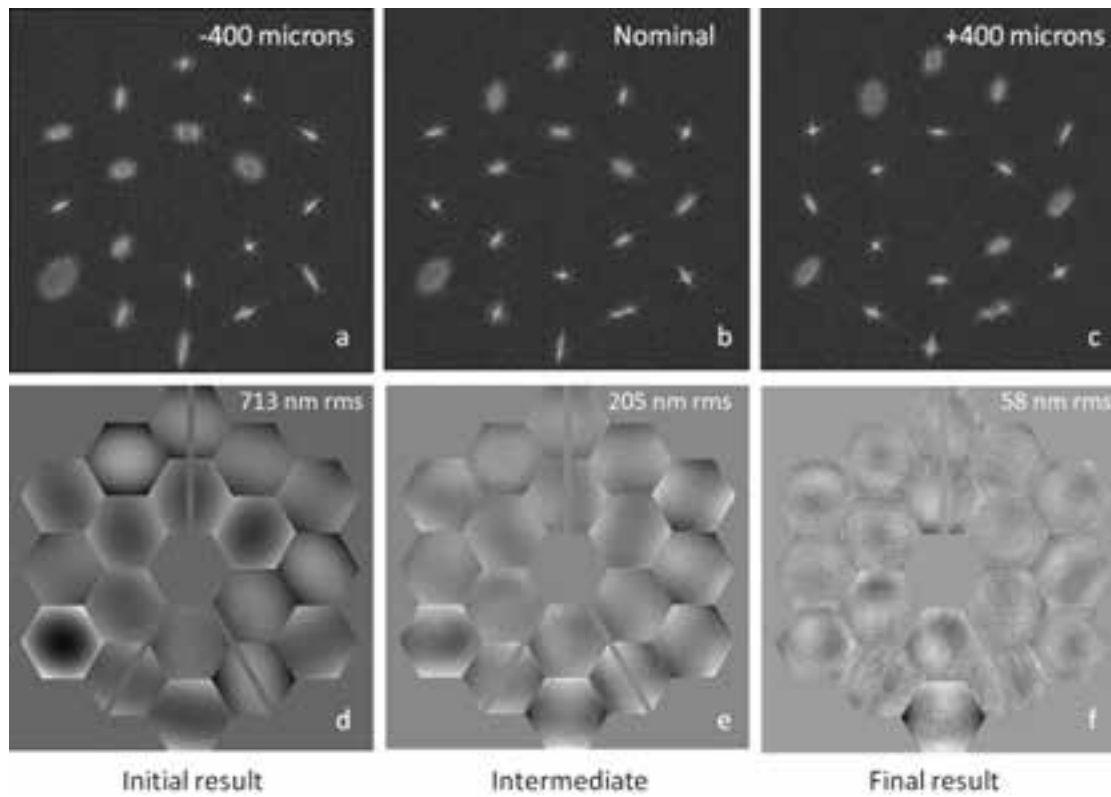
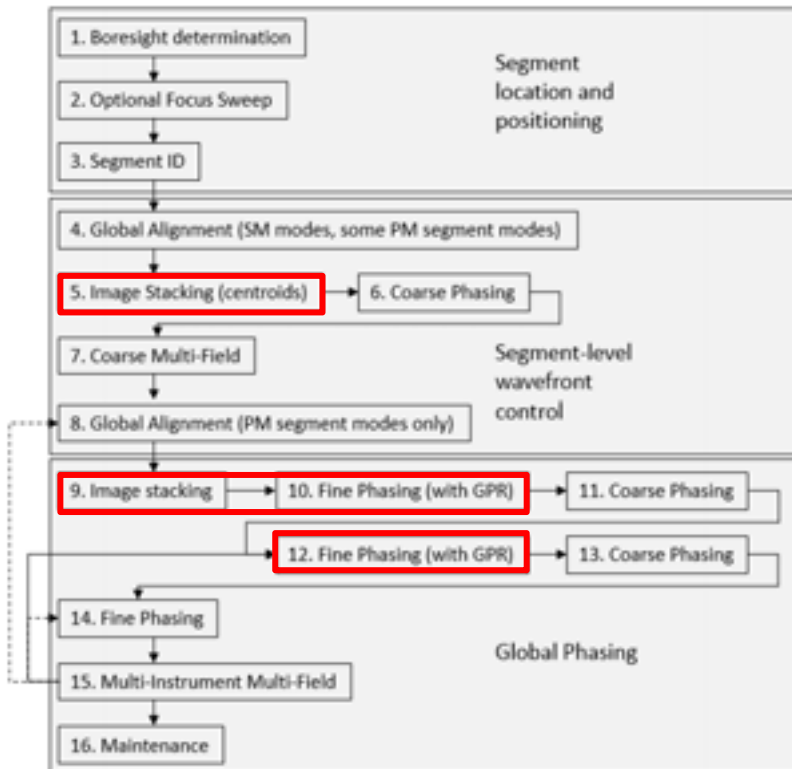


Image array

Resultant wavefront

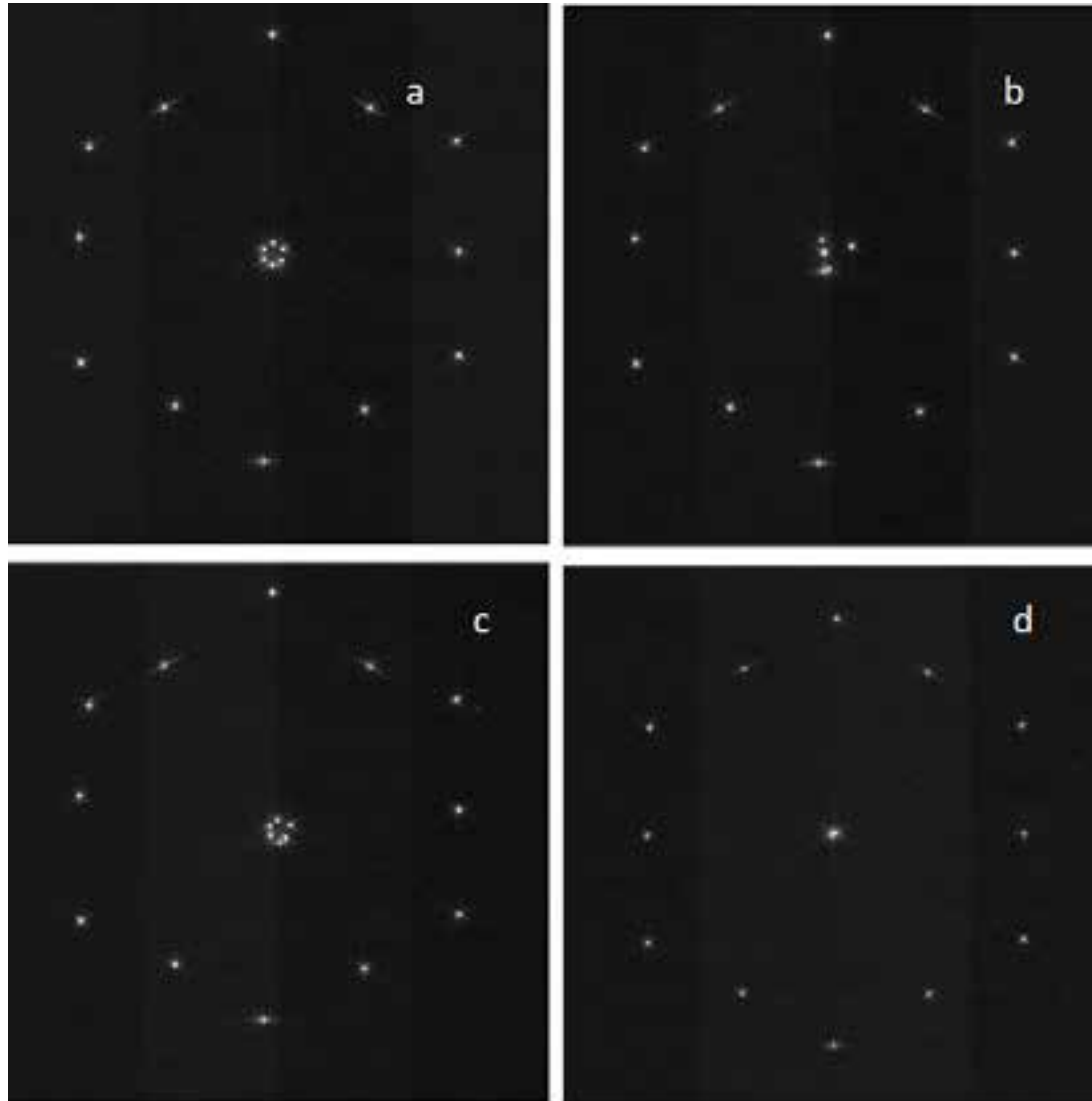


# Image Stacking



- Image stacking is needed in preparation for coarse phasing (performed 3 times).
- The first stacking is based on centroids
- Subsequent stacking uses a Geometrical Phase Retrieval algorithm (GPR).
- In each case, moves must be made while conserving the fine range of the actuators.
- Guiding must be active during the stacking process

# Image Stacking

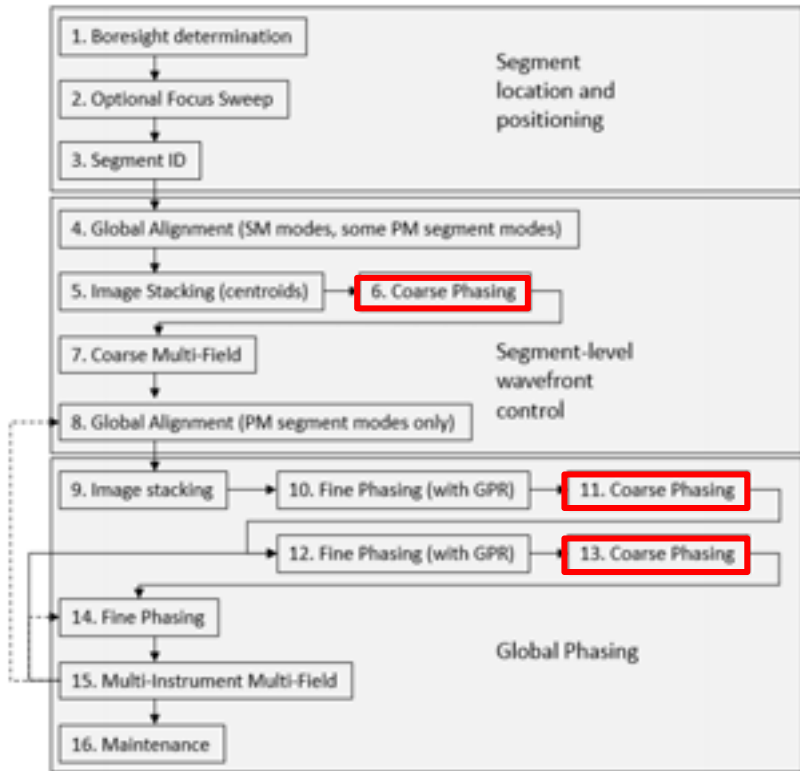


Stacking based on centroids.

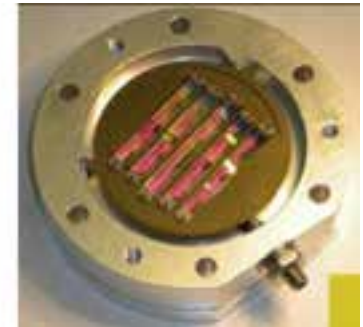
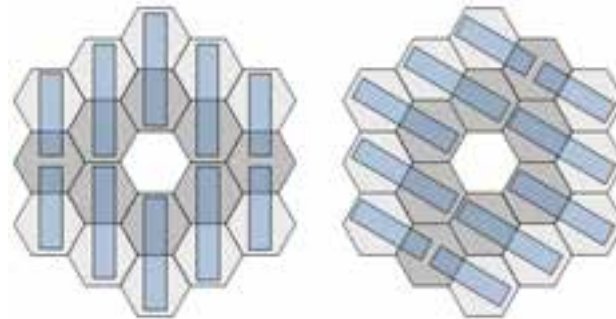
- Guiding is performed on a spot near the perimeter.
- (a) Six spots are moved near the center. Their positions are measured.
- (b) Spots are commanded to center and fine mechanisms are recentered. They are measured and moved back to the small array.
- (c) Their positions are measured.
- (d) The spots are stacked.

This is repeated for the remaining segments in groups of 6. Guiding is transferred to the stacked segments during the last group.

# Coarse Phasing



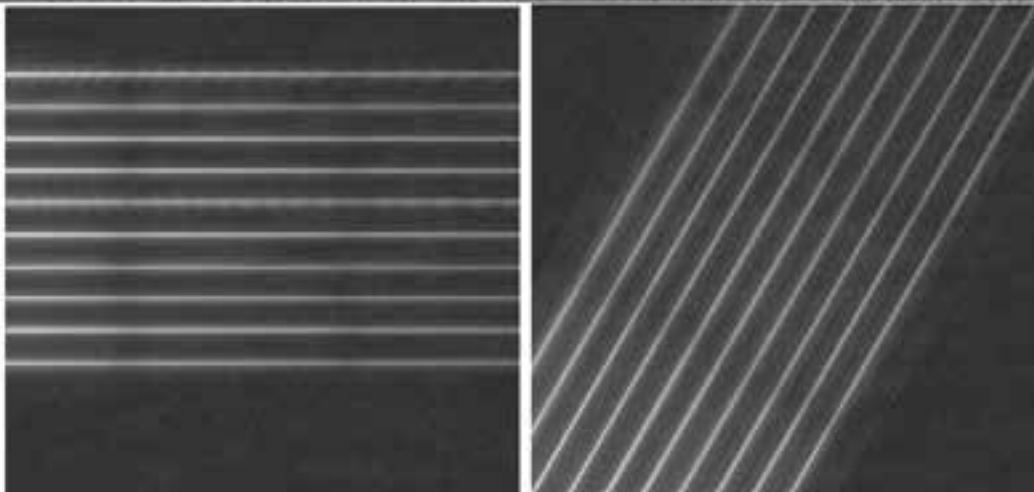
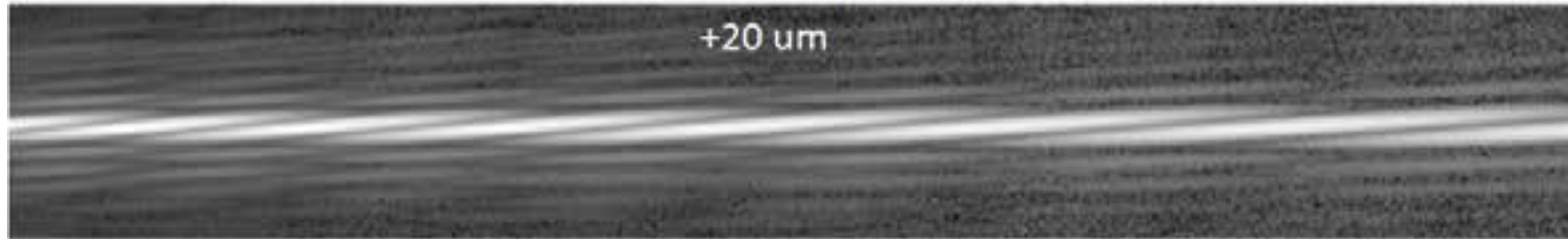
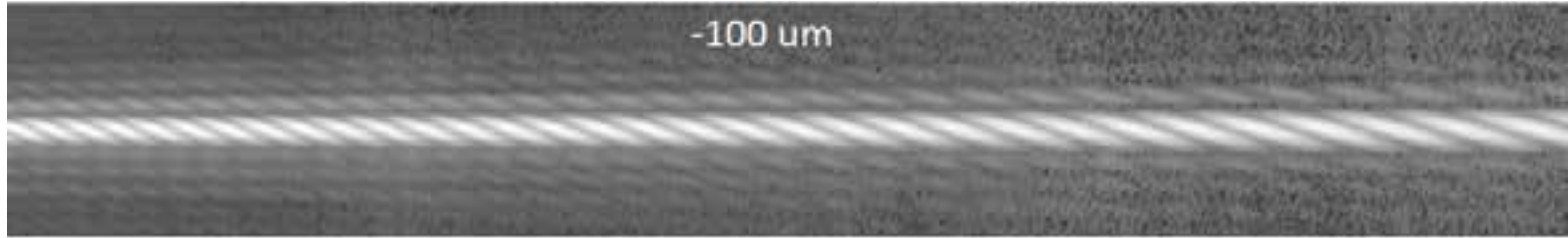
- Coarse Phasing is performed 3 times.
- Goal is to reduce PMSA piston errors from >200 microns to ~0.3 microns
- Uses Dispersed Fringe Sensing (DFS) technology



# Coarse Phasing

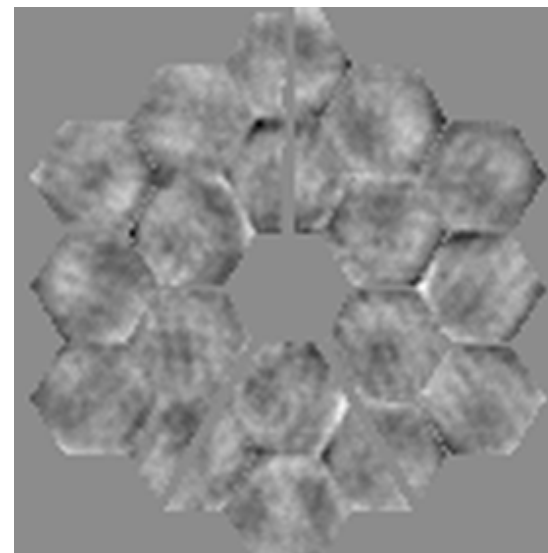
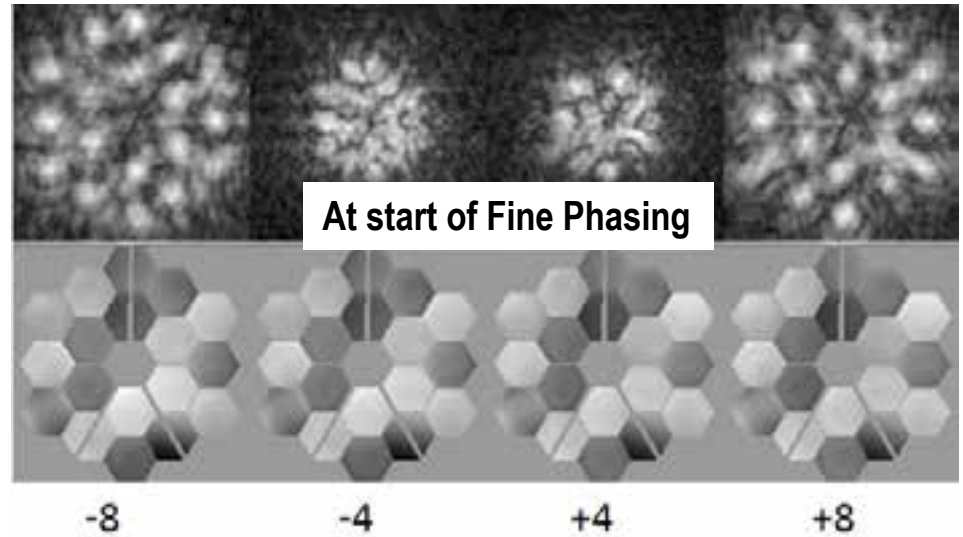
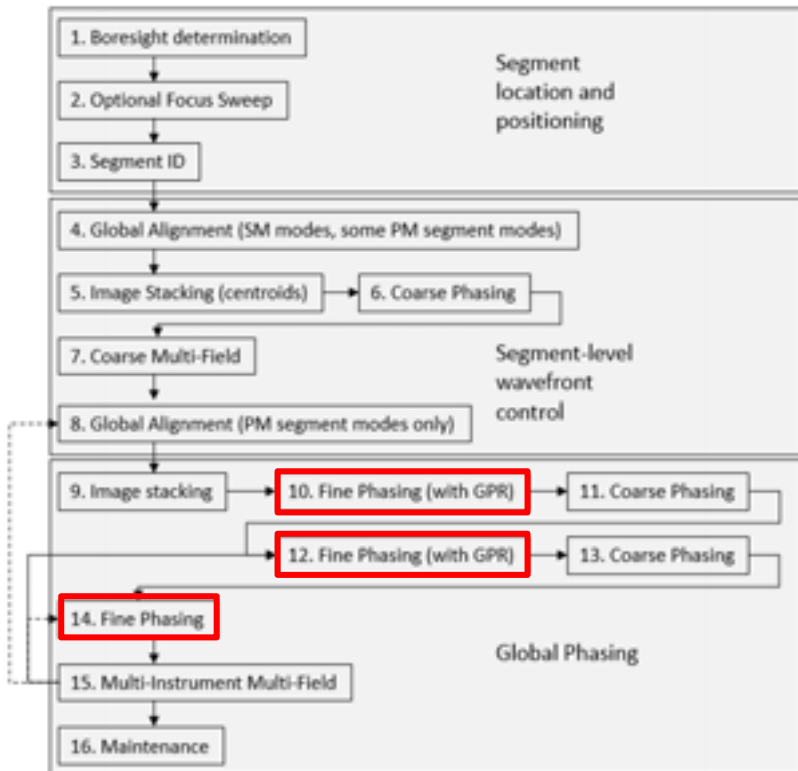


Sample spectra



Full-frame NIRCcam Images

# Fine Phasing



- Weak lenses are placed in the filter elements of NIRCcam to create defocused images
- A Phase Retrieval algorithm is applied to calculate the wavefront
- Corrections are applied to completely phase the telescope, at a single field point

# Multi-instrument Multi-field



- Uses the SM to create slightly defocused images a ~ 30 field points on multiple science instruments.
- Phase retrieval algorithm calculates wavefront from these images. This is challenging because...
  - The images are generally broad-band
  - The image are often undersampled
  - On NIRSpec, the micro-shutter array and an internal stop are in the beam path
  - Minimal defocus is used, due to the need to guide during exposures
- Monochromatic, fully-sampled weak lens images (as used in Fine Phasing) generate the fine details in the wavefront.
- The other Phase Retrieval algorithms use this as a starting point.
- A pseudo-linear model is used to generate an optimal solution.